

Our collaboration is not only work...



Challex



Chateau in the Loire



Sicily



1988: U. Becker 50th Birthday celebration at CERN, Ting's office



Visit of Chuck Vest, Paul Gray, and the MIT Corporation to L3



Christoph Paus at L3 reception with Becker



SCIENTIFIC AMERICAN

The smart genes that control cell differentiation.

What happens when galaxies collide.

Can antichaos explain the origin of species?

AUGUST 1991
\$3.50

~ 300 Papers

Results:

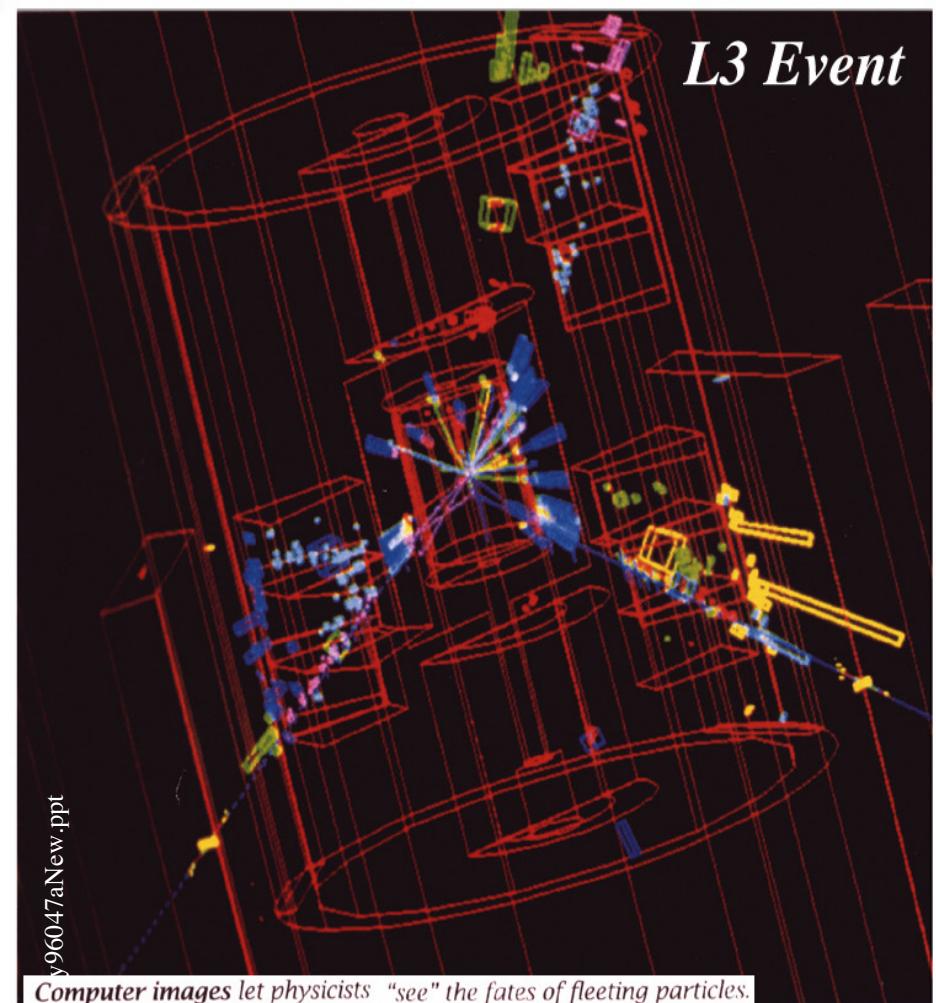
3 kinds of electrons (e, μ, τ)

6 families of quarks

Electrons and quarks have no size

Their radii are $R < 10^{-17}$ cm

The results are in excellent agreement with Electroweak Theory



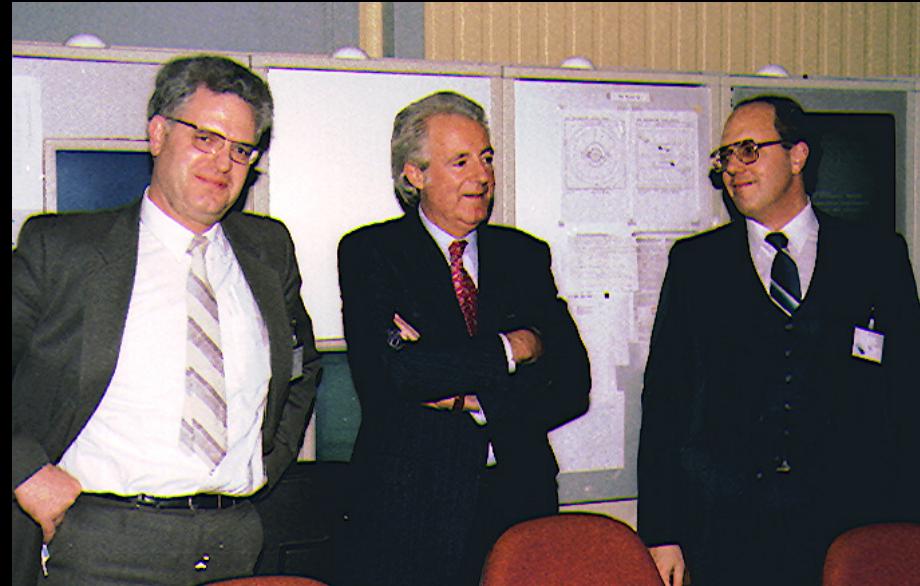
Computer images let physicists "see" the fates of fleeting particles.

Discussions of L* experiment at SSC

L. Okun



K. Lubelsmeyer



H. Newman

M. Vivargent H. Hofer F. Wittgenstein



Z P Zheng



1994, Sixth Experiment with Ulrich Becker AMS in Space

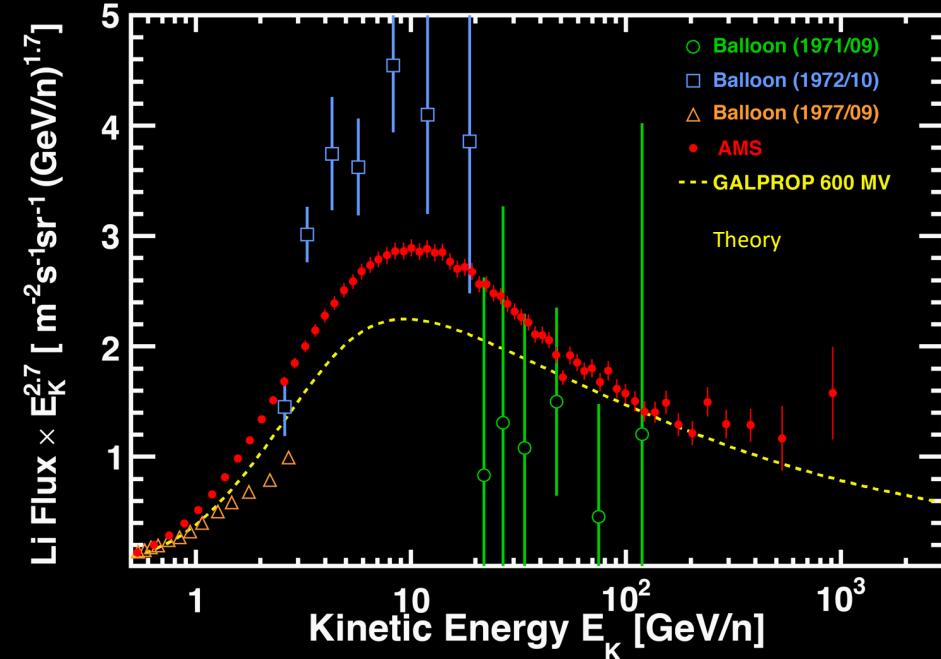
Space is the ultimate laboratory. It provides the highest energy particles.

The Space Station is a unique platform to support the weight and to provide the power for a precision, long duration experiment, AMS.

In the past hundred years, measurements of charged cosmic rays by balloons and satellites have typically had 30% to 50% accuracy.

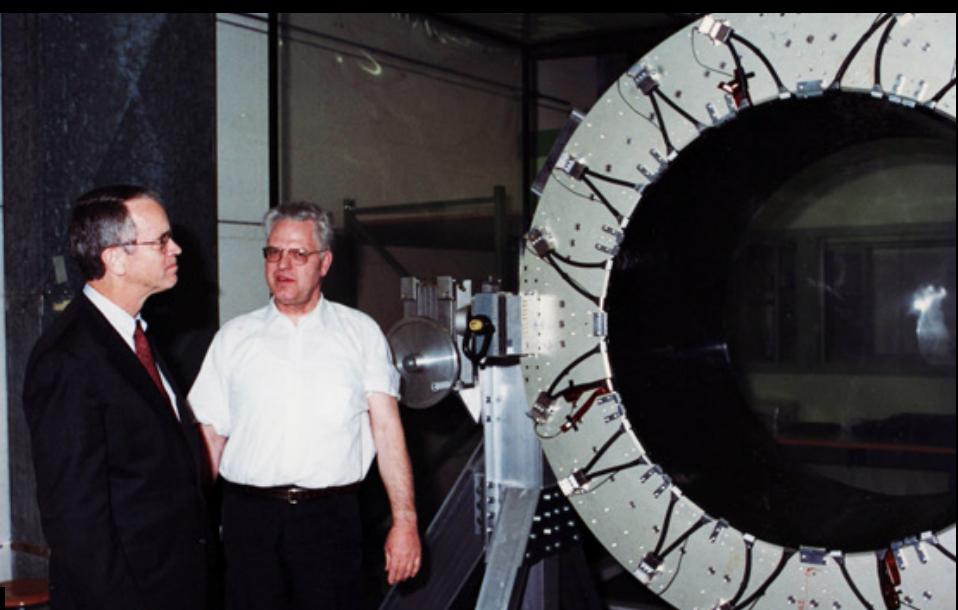
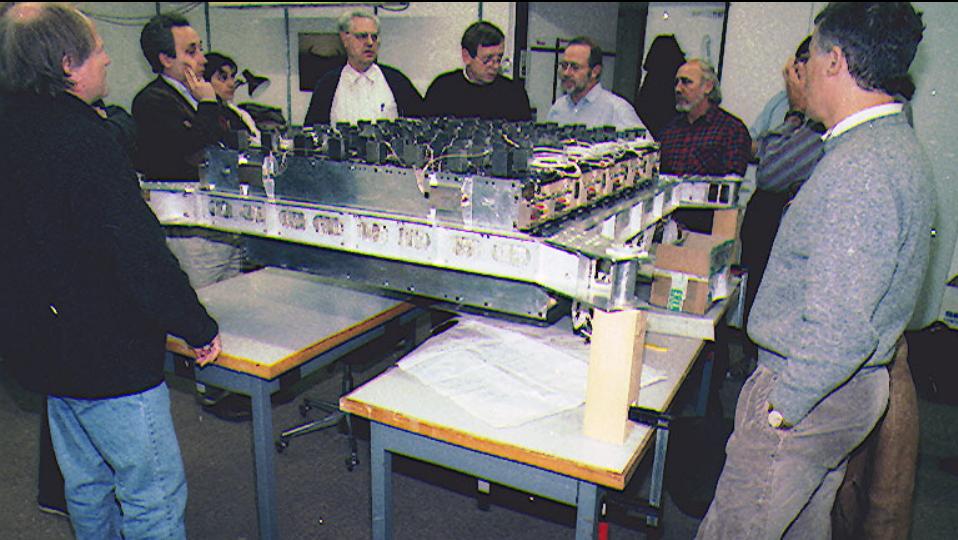
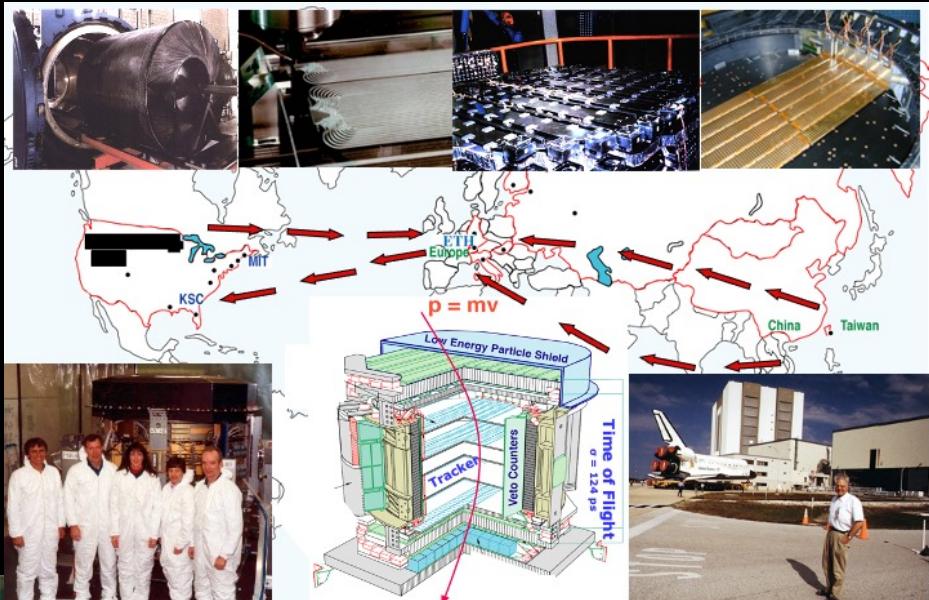
AMS is providing cosmic ray information with ~1% accuracy.

The improvement in accuracy is providing new insights about the cosmos.



AMS-01 on the space shuttle

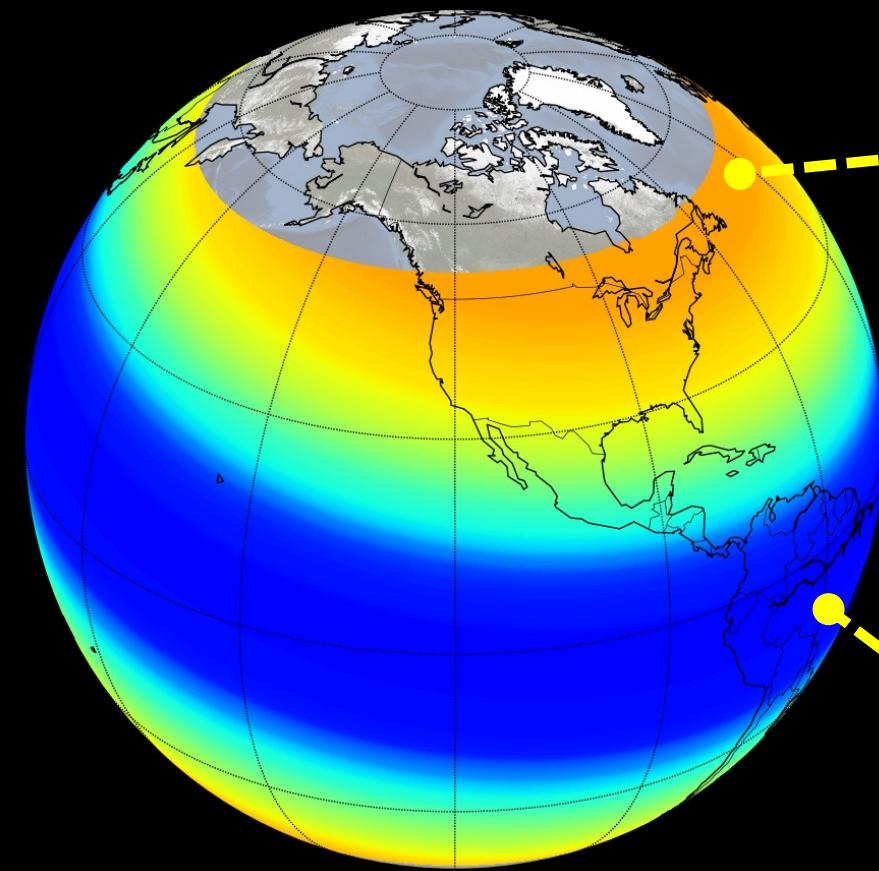
Approval: April 1995, Assembly: December 1997, Flight: 10 days in June 1998



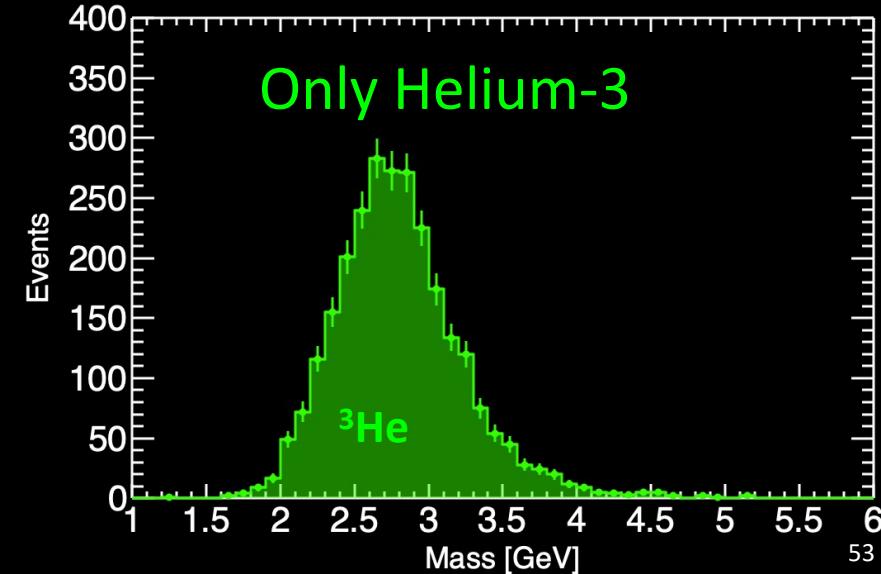
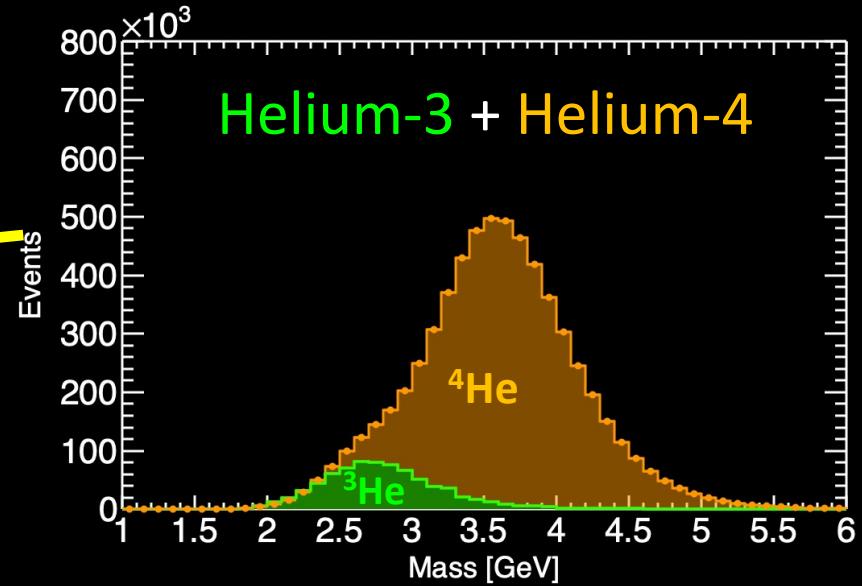
Helium in Low Earth Orbit

In the energy range 100 MeV to 4000 MeV
over a large region, $-35^\circ < \Theta_M < +35^\circ$, only helium-3 isotope exists.

Helium



10^{-1} 10^2
Flux [Particle/($m^2 \times sr \times sec$)]



May 2011

CERN COURIER
January/February 2010 cerncourier.com
Reporting on international high-energy physics

AMS KEEPS ITS COOL

Atomki anomaly rekindled • Tackling the flavour puzzle • Voyage to the neutrino mass



2019 upgrade

The Space Station's Crown Jewel

A fancy cosmic-ray detector, the Alpha Magnetic Spectrometer, is about to scan the cosmos for dark matter, antimatter and more

By George Musser, staff editor

THE WORLD'S MOST ADVANCED COSMIC-RAY DETECTOR TOOK 16 YEARS AND \$2 billion to build, and it's finally headed to the International Space Station. It will be installed on the space station next year, by the end of 2012, while it simply did not have room in its schedule to launch the instrument privately. Seeing it took a lobbying campaign by physicists and intervention by the U.S. Congress, which insisted that the instrument be sent on the shuttle Endeavour, scheduled to take off April 19 to express the purpose of developing the Alpha Magnetic Spectrometer (AMS) at the International Space Station.

Compared to other particle and atomic nodes that fly up in balloons or come from particle accelerators, the AMS is a giant leap forward. No one has who knows what—the last category naturally being of greatest interest, and the main impetus for this new instrument. Dark matter is one of those possible mystery particles. Clumps of it could pull galaxies together, and it may be the source of the universe's missing sectors of light. Some physicists speculate that the planet might be peppered with the odd atom or two from distant galaxies made out of matter but of no mass.

The spectrometer is built to find it, and it's built to do it in the most efficient way, which means it's built to confirm. No other instrument has the combination of detectors that can tease out all the properties of a particle: mass, velocity, type, electric charge.

In closer processor is the PAMELA instrument, launched by a European consortium in 2002. The AMS is built to surpass it in every way, and it's built to do it in a much smaller package because it lacks the ability to distinguish a low-mass antiproton, such as a positron, from high-mass ordinary particle with the same electric charge, such as a proton.

The AMS instrument is monitored by the standard of the space program, with a mass of seven tons and a power requirement of 200 watts. The PAMELA instrument has a mass of 2,400 watts, in a compact symbolic way. At the space station the crew is justly worth others' interests. The scientists who built the AMS will be there for power and orbits, and the spectators will be there to look at the results. And the AMS will be there for the long haul. Its 15-year mission will do world-class research. As CERN's Lynne Hoddeson charter the depths of nature will be done from orbit. The Alpha Magnetic Spectrometer will do some from orbit.

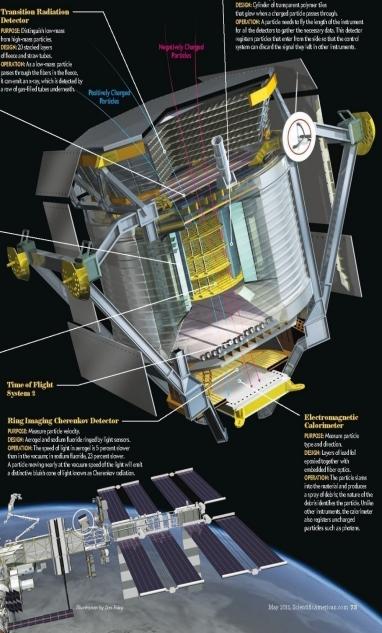
SCIENTIFIC AMERICAN ONLINE
For more on the AMS, visit our website for the latest news on the instrument's work.
ScientificAmerican.com/sc0511ams

72 Scientific American, May 2011

Scientific American 2011

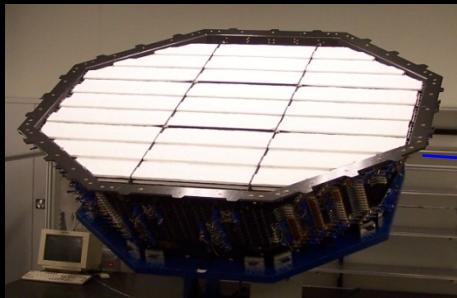


AMS-02 on Space Station

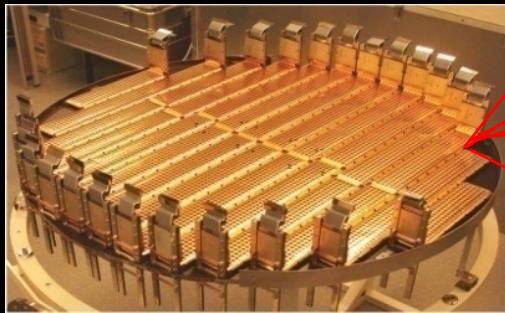


AMS is a space version of a precision detector used in accelerators

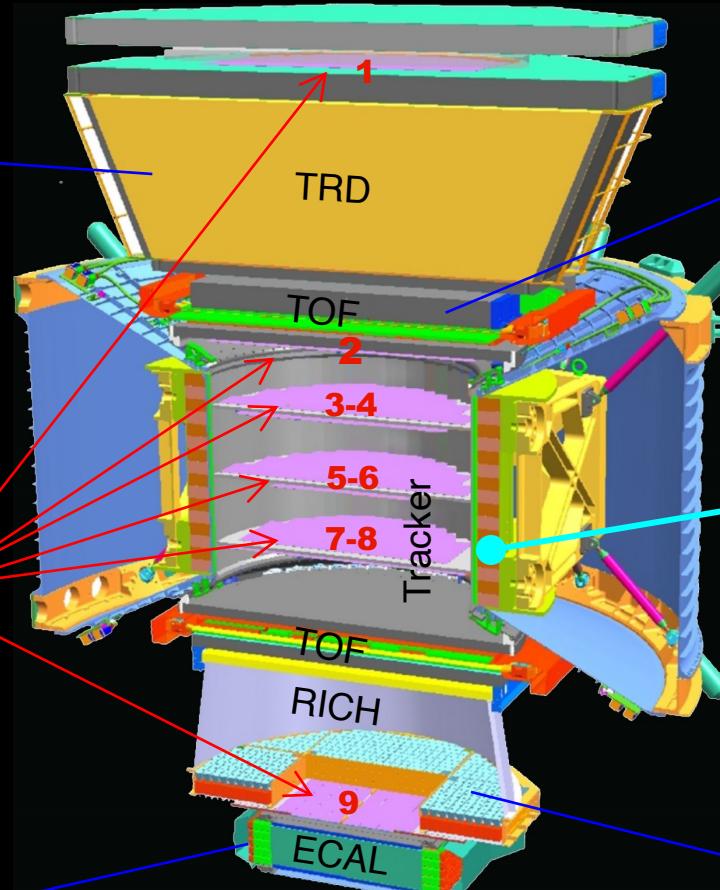
Transition Radiation Detector (TRD)



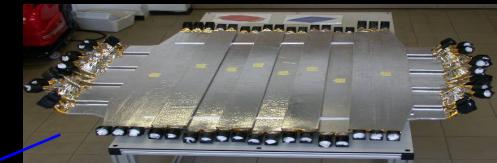
Silicon Tracker



Electromagnetic Calorimeter (ECAL)



Time of Flight Detector (TOF)



Magnet

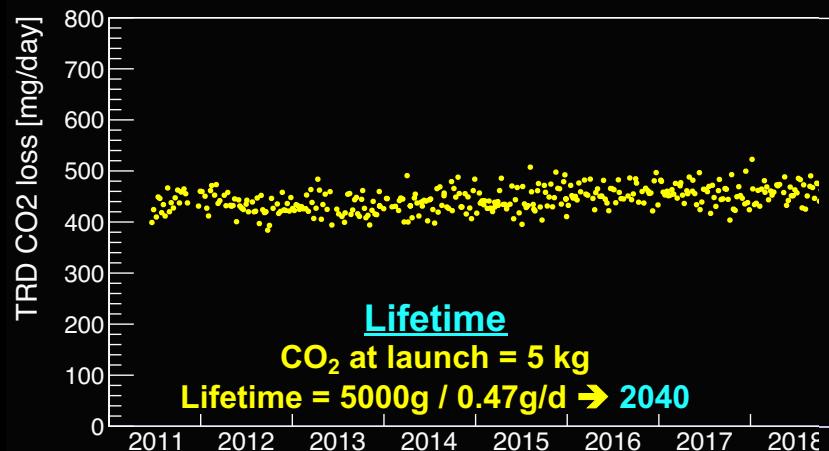
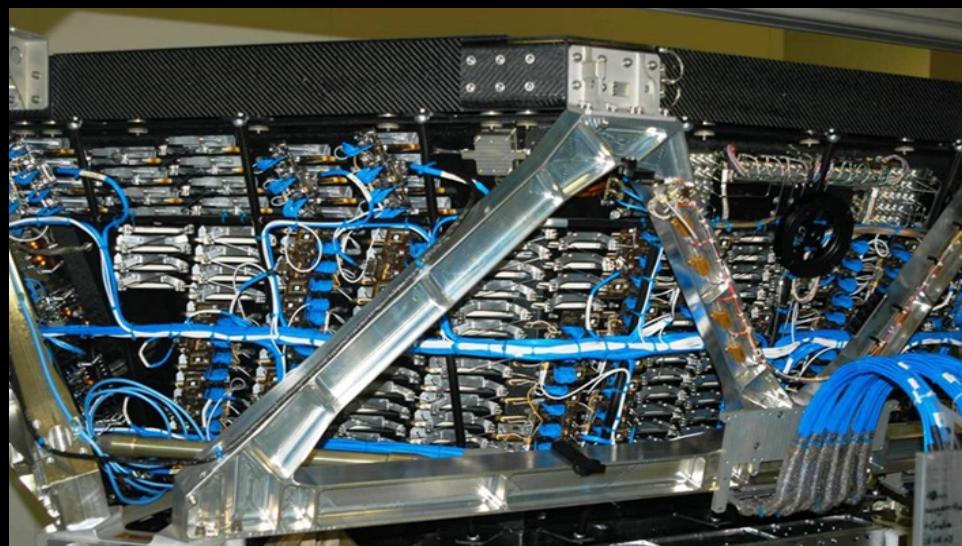


Ring Imaging Cherenkov (RICH)

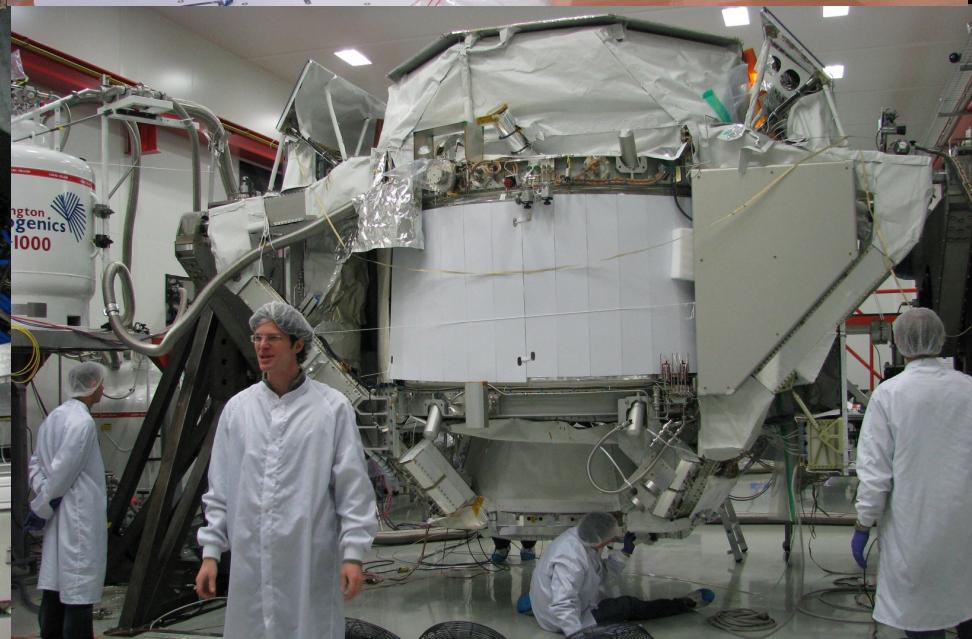


**300,000 electronic channels,
650 fast microprocessors
5m x 4m x 3m
7.5 tons**

TRD Gas system: U. Becker, J. Burger, P. Fisher



The Beckers in AMS



Robert Becker has made important contributions to AMS

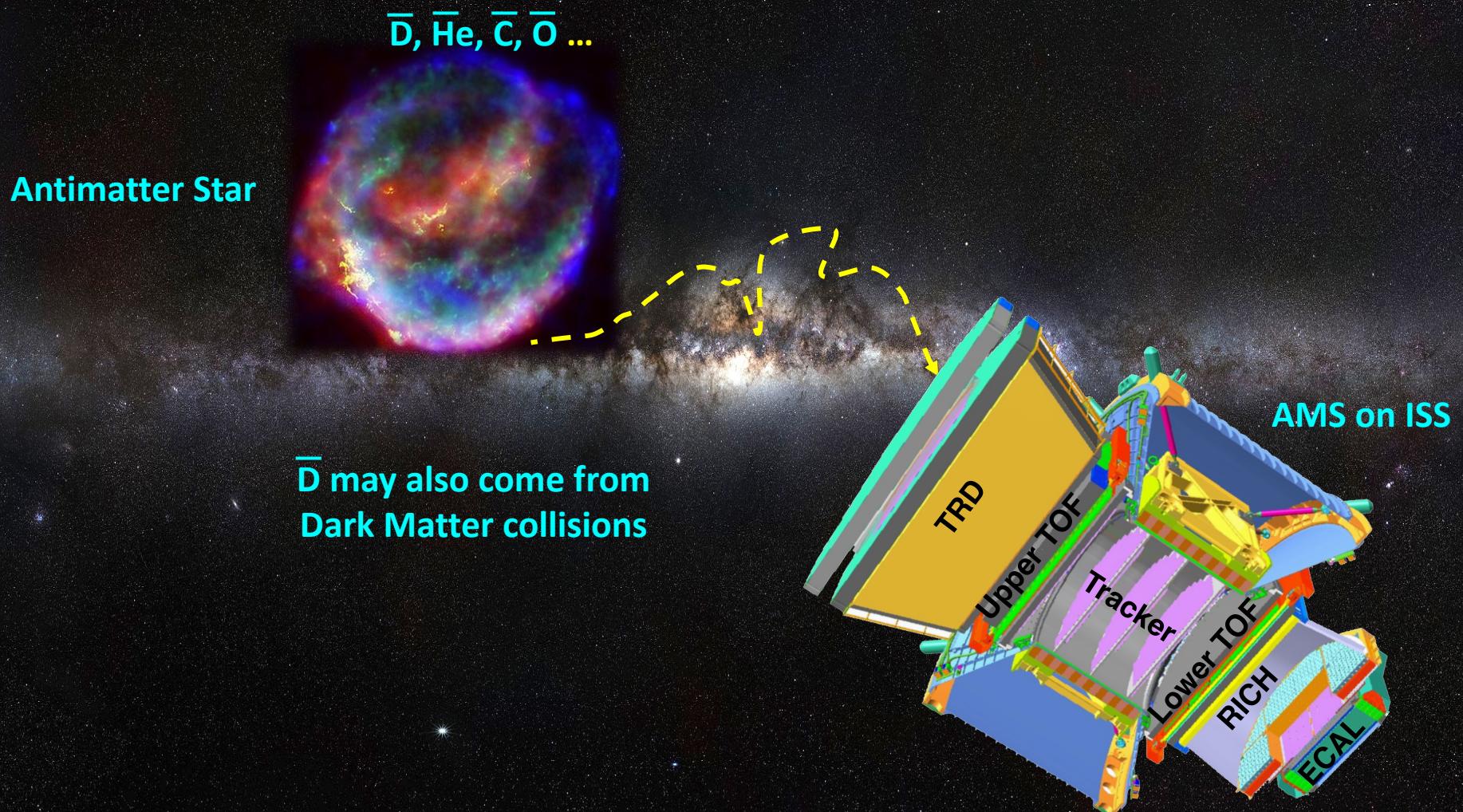
AMS-02 Publications in P.R.L. with U.J. Becker

- 1) Phys. Rev. Lett. 110, 141102 (2013). Editors' Suggestion. Viewpoint in *Physics*.
Highlight of 2013. Ten-Year retrospective.
- 2) Phys. Rev. Lett. 113, 121101 (2014). Editors' Suggestion
- 3) Phys. Rev. Lett. 113, 121102 (2014). Editors' Suggestion
- 4) Phys. Rev. Lett. 113, 221102 (2014).
- 5) Phys. Rev. Lett. 114, 171103 (2015). Editors' Suggestion
- 6) Phys. Rev. Lett. 115, 211101 (2015). Editors' Suggestion
- 7) Phys. Rev. Lett. 117, 091103 (2016).
- 8) Phys. Rev. Lett. 117, 231102 (2016). Editors' Suggestion
- 9) Phys. Rev. Lett. 119, 251101 (2017).
- 10) Phys. Rev. Lett. 120, 021101 (2018). Editors' Suggestion. Featured in *Physics*.
- 11) Phys. Rev. Lett. 121, 051101 (2018).
- 12) Phys. Rev. Lett. 121, 051102 (2018). Editors' Suggestion
- 13) Phys. Rev. Lett. 121, 051103 (2018).
- 14) Phys. Rev. Lett. 122, 041102 (2019). Editor's Suggestion
- 15) Phys. Rev. Lett. 122, 101101 (2019).
- 16) Phys. Rev. Lett. 123, 181102 (2019). Editors' Suggestion
- 17) Phys. Rev. Lett. 124, 211102 (2020). Editors' Suggestion. Featured in *Physics*.

Heavy Antimatter in the Cosmos

Matter is defined by its mass M and charge Z .

Antimatter has the same mass M but opposite charge $-Z$.



AMS is a unique antimatter spectrometer in space

Antimatter events

